The NOvA Experiment

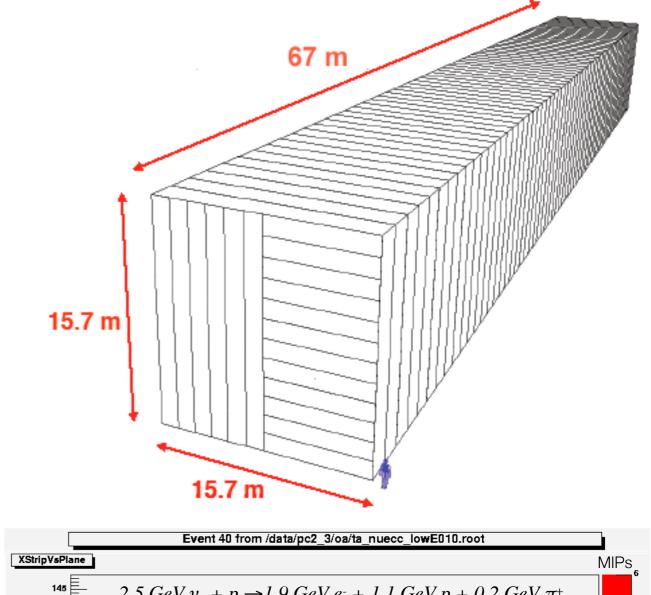
Mark Messier Indiana University

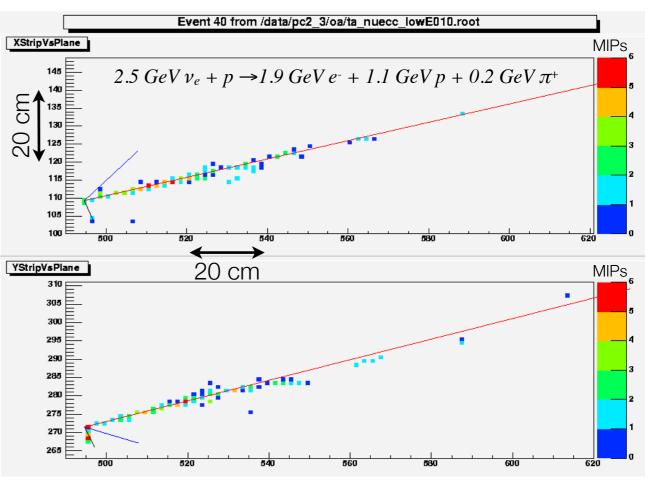
for the NOvA collaboration 181 scientists and engineers, 26 institutions http://www-nova.fnal.gov

3rd Workshop on Physics with a high intensity proton source June 5-6, 2008 Fermilab

The NOvA Experiment

- NOvA is a second generation experiment on the NuMI beamline which is optimized for the detection of v_µ→v_e and v̄_µ→v̄_e oscillations
- NOvA is:
 - An upgrade of the NuMI beam intensity from 400 kW to 700 kW
 - A 15 kt "totally active" tracking liquid scintillator calorimeter sited 14 mrad off the NuMI beam axis at a distance of 810 km
 - A 215 ton near detector identical to the far detector sited 14 mrad off the NuMI beam axis at a distance of 1 km





As the first chapter in the study of neutrino oscillations comes to an end, a new chapter begins. The great progress in neutrino physics over the last few decades raises new questions and provides opportunities for major discoveries. Among the compelling issues today:

- What is the value of θ₁₃, the mixing angle between first- and third-generation neutrinos for which, so far, experiments have only established limits? Determining the size of θ₁₃ has critical importance not only because it is a fundamental parameter, but because its value will determine the tactics to best address many other questions in neutrino physics.
- 2) Do neutrino oscillations violate CP? If so, how can neutrino CP violation drive a matter-antimatter asymmetry among leptons in the early universe (leptogenesis)? What is the value of the CP violating phase, which is so far completely unknown? Is CP violation among neutrinos related to CP violation in the quark sector?
- 3) What are the relative masses of the three known neutrinos? Are they "normal," analogous to the quark sector, (m₃>m₂>m₁) or do they have a so-called "inverted" hierarchy (m₂>m₁>m₃)? Oscillation studies currently allow either ordering. The ordering has important consequences for interpreting the results of neutrinoless double beta decay experiments and for understanding the origin and pattern of masses in a more fundamental way, restricting possible theoretical models.
- 4) Is θ₂₃ maximal (45 degrees)? if so, why? Will the pattern of neutrino mixing provide insights regarding unification of the fundamental forces? Will it indicate new symmetries or new selection rules?
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- 6) What can we learn from observation of the intense flux of neutrinos from a supernova within our galaxy? Can we observe the neutrino remnants of all supernovae that have occurred since the beginning of time?
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- 8) What can neutrinos tell us about new physics beyond the Standard Model, dark energy, extra dimensions? Do sterile neutrinos exist?

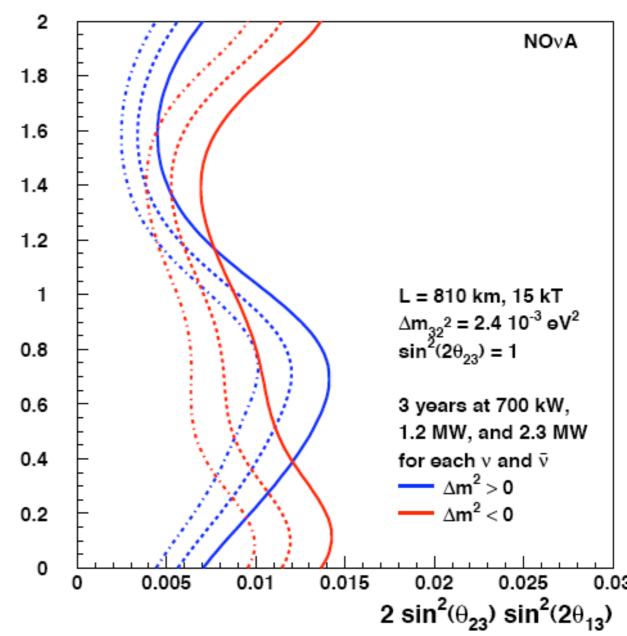
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1) What is the value of θ_{13} ?

90% CL Sensitivity to $\sin^2(2\theta_{13}) \neq 0$

δ (π)



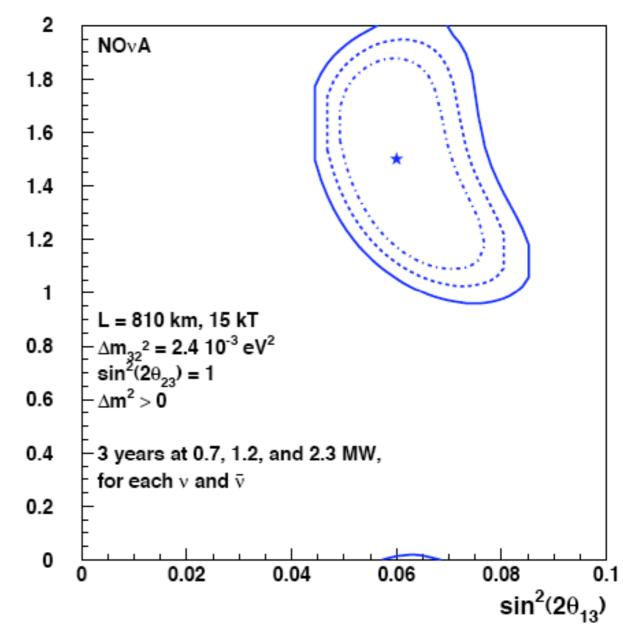
NOvA searches for electron neutrino appearance down to ~0.01 at 90% CL

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2) Do neutrino oscillations violate CP?

1 σ Contours for Starred Point

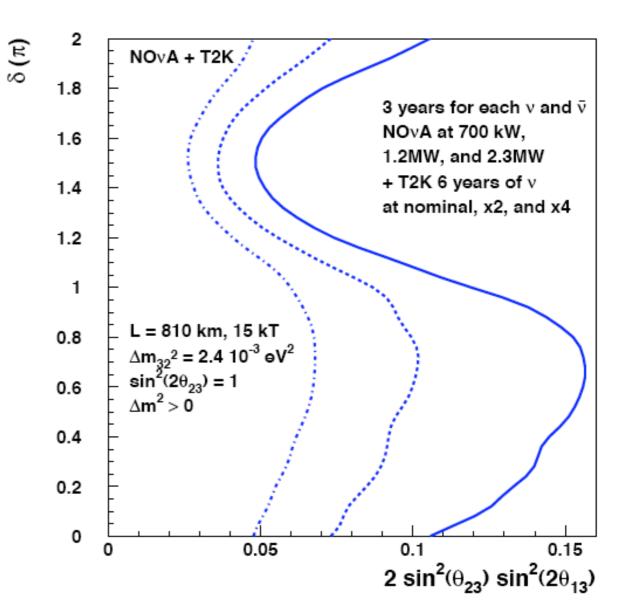


NOvA provides the first look into the CPV parameter space

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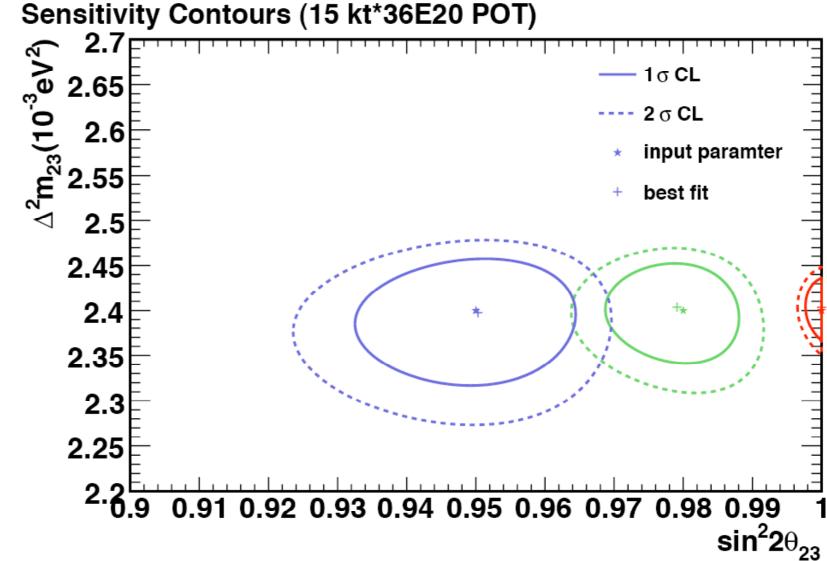


NOvA's long baseline makes it sensitive to the mass ordering

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4) Is θ_{23} maximal?

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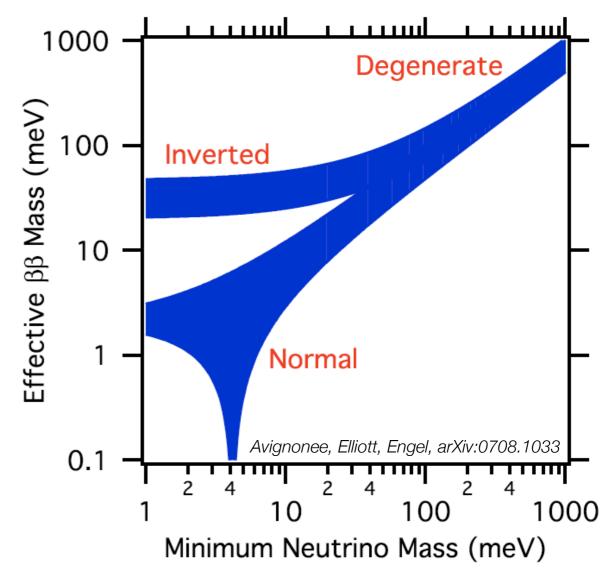


Because of its excellent energy resolution NOvA can make ~1% measurements of muon neutrino disappearance using quasi-elastic channel

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5) Are neutrinos their own antiparticles?

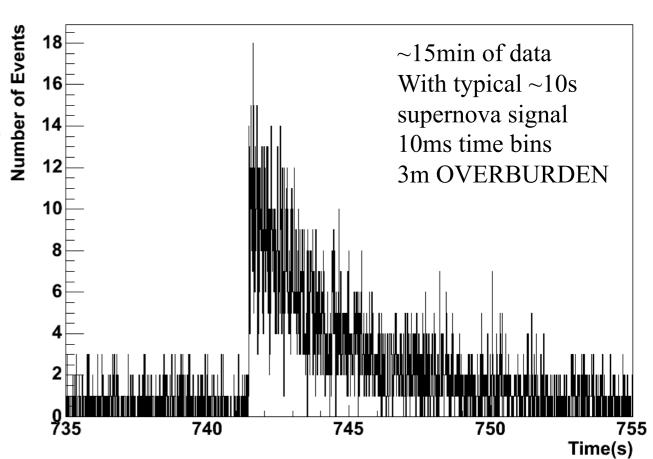


If NOvA establishes inverted hierarchy and next generation of $0\nu\beta\beta$ experiments see nothing, then it is very likely that neutrinos are Dirac particles

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6) ...supernova within our galaxy?



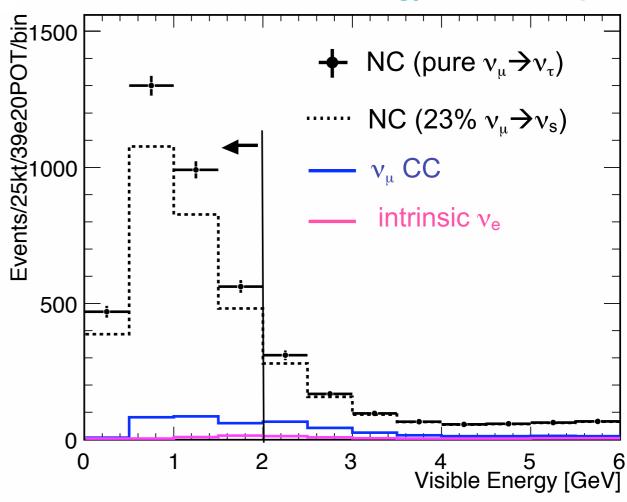
NOvA would see burst of 5000 events for a supernova at the center of the galaxy

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8) ...beyond the Standard Model...Do sterile neutrinos exist?

Reconstructed visible energy for NC sample



NOvA's granularity allows for clean neutral-current measurements facilitating searches for sterile neutrinos

Schedule

- Despite FY08 omnibus work continues
 - NOvA passed DOE Office of Science CD2/3a in Oct. 2007.
 - Passed External Independent Review (EIR) in Nov. 2007
 - FY08 Omnibus required repeats. Passed DOE Office of Science CD2 review again in April 2008. EIR scheduled for June 19.
 - Expect CD3a in Feb. 2009, CD3b in Sept. 2009
 - Were it not for the omnibus, construction would have started 3 months ago
- Cost & schedule adjusted to include FY08 omnibus and anticipated 4 month continuing resolution in FY09
 - Start of construction April 2009
 - First 2.5 kt taking data August 2012
 - Detector complete January 2014
- NOvA construction schedule is driven by funding profile. We know what we want to build and we could build it faster.

NOvA and the P5 roadmap

NOvA is the foundation of the US accelerator neutrino program

- It addresses 7 of the 8 physics questions called out by P5 as the focus of the neutrino program over the next decade
- Among the next generation experiments, NOvA uniquely provides information on the mass hierarchy and CP phase
- NOvA provides the incentive and continuity to increase the NuMI beam power from 400 to 700 kW and ultimately to 2.3 MW
- Ensures a robust future program.

 | 2.3 MW | V beam to DUSEL? |
 | Mt H2O? | 100+ kt LgAr? |
 | 700 kW V beam to DUSEL? |
 | 5 kt LgAr? |
 | NOVA